

28) Bose-Einstein-Kondensation in d Dimensionen

a) Zustandsdichte:

$$\sum_p \rightarrow \frac{V}{h^d} \int d^d p = \frac{V}{h^d} S_d \int dp p^{d-1} = \int dE g(E)$$

$$dp p^{d-1} d\Omega_d \quad p = \left(\frac{E}{a}\right)^{1/r} \Rightarrow dp = \frac{1}{a^{1/r}} \frac{1}{r} E^{\frac{1}{r}-1} dE$$

$$p^{d-1} = \left(\frac{E}{a}\right)^{\frac{d-1}{r}}$$

$$dp p^{d-1} = dE \frac{E^{\frac{d}{r}-1}}{r a^{d/r}}$$

$$\Rightarrow g(E) = \frac{V}{h^d} \underbrace{\frac{S_d}{r a^{d/r}}}_{=: g_0} E^{\frac{d}{r}-1}$$

mit $S_d = \frac{2\pi^{d/2}}{\Gamma(d/2)}$, vgl. Bsp. 12.

b) Grosskanon. Potential:

$$Z_{GK} = \prod_p \frac{1}{1 - e^{-\beta(E_p - \mu)}} \Rightarrow \beta\Omega = -\ln Z_{GK} = \sum_p \ln(1 - e^{-\beta(E_p - \mu)})$$

$$\approx \underbrace{\ln(1-z)}_{\sim \text{Grundzustand}} + \underbrace{\int dE g(E) \ln(1 - z e^{-\beta E})}_{\sim \text{Gaussezustand}}$$

2. Term:

$$g_0 \int dE E^{\frac{d}{r}-1} \ln(1 - z e^{-\beta E})$$

$$\stackrel{\text{P.I.}}{=} g_0 \frac{r}{d} E^{\frac{d}{r}} \ln(1 - z e^{-\beta E}) \Big|_0^\infty - g_0 \frac{r}{d} \int_0^\infty dE E^{\frac{d}{r}} \frac{\beta z e^{-\beta E}}{1 - z e^{-\beta E}}$$

$$= -g_0 \beta \frac{r}{d} \int_0^\infty dE E^{\frac{d}{r}} \frac{1}{z^{-1} e^{\beta E} - 1} = -g_0 \frac{1}{\beta^{d/r}} \frac{r}{d} \int_0^\infty dx \frac{x^{d/r}}{z^{-1} e^x - 1}$$

$$\Rightarrow \beta\Omega = \ln(1-z) - g_0 \frac{\Gamma(d/r)}{\beta^{d/r}} g_{d/r+1}(z)$$

$\frac{\Gamma(d/r+1) g_{d/r+1}(z)}{d/r \Gamma(d/r)}$

Innere Energie:

$$U = \sum_p \frac{E_p}{e^{\beta(E_p - \mu)} - 1} = \frac{\partial(\beta\Omega)}{\partial\beta} \Big|_{z, V} = g_0 \frac{d}{r} \frac{\Gamma(d/r)}{\beta^{d/r+1}} g_{d/r+1}(z)$$

$$= \frac{d}{r\beta} \left(\ln(1-z) - \underbrace{\beta\Omega}_{-PV} \right) = \frac{d}{r} \left(PV + \frac{1}{\beta} \ln(1-z) \right) \xrightarrow[V \rightarrow \infty]{\text{da } 1-z \sim \frac{1}{V}, \text{ siehe unten.}} \frac{d}{r} PV$$

Bsp: d=3 Dimensionen:

-) nichtrelativistisch (r=2): $U = \frac{3}{2} PV$
 -) ultrarelativistisch (r=1): $U = 3PV$
- vgl. Bsp. 26, 27 für Fermionen

c) Mittlere Teilchendichte:

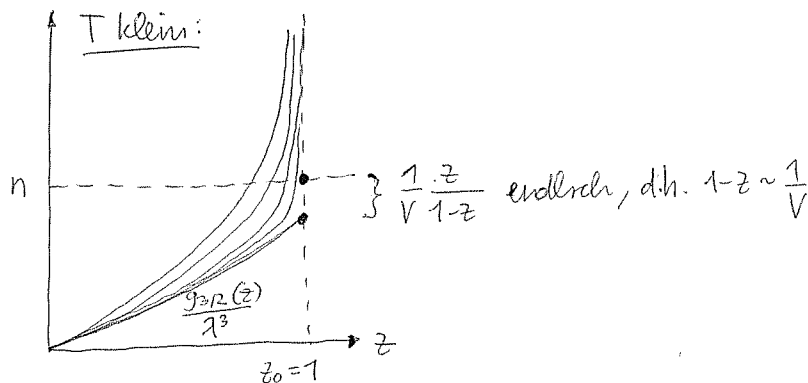
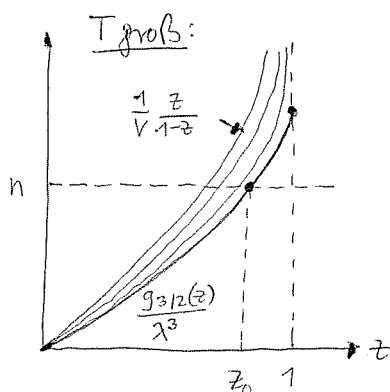
$$\begin{aligned}\langle N \rangle &= \sum_r \frac{1}{e^{\beta(\epsilon_r - \mu)} - 1} = - \frac{\partial \Omega}{\partial \mu} = - \frac{\partial \Omega}{\partial z} \frac{\partial z}{\partial \mu} = - \beta z \frac{\partial \Omega}{\partial z} = \\ &= - \beta z \left[- \frac{1}{\beta} \frac{1}{1-z} - g_0 \frac{r}{d} \frac{\Gamma(d/r+1)}{\beta^{d/r+1}} \underbrace{\frac{d}{dz} g_{d/r+1}(z)} \right] \\ &= \frac{z}{1-z} + g_0 \frac{r}{d} \frac{\Gamma(d/r+1)}{\beta^{d/r}} g_{d/r}(z) \quad \text{aus } g_\alpha(z) = \sum_{n=1}^{\infty} \frac{z^n}{n^\alpha} : z \frac{d}{dz} g_\alpha(z) = g_{\alpha-1}(z) \\ &= \frac{z}{1-z} + g_0 \frac{\Gamma(d/r)}{\beta^{d/r}} g_{d/r}(z)\end{aligned}$$

$$\Rightarrow n = \frac{\langle N \rangle}{V} = \underbrace{\frac{1}{V} \frac{z}{1-z}}_{\sim \frac{1}{V}} + \underbrace{\frac{S_d}{r h^d} \frac{\Gamma(d/r)}{(\alpha \beta)^{d/r}} g_{d/r}(z)}_{\text{hängt von Temperatur ab.}} \quad \dots \quad \begin{array}{l} \text{für } n, \beta = \text{const:} \\ \text{Bestimmungsgleichung für } z, \\ z \in [0, 1] \end{array}$$

für $r=2, a = \frac{1}{2m} \Rightarrow \frac{1}{\lambda^d}$

Uns interessiert $V \rightarrow \infty$. z.B. für $d=3, r=2, a = \frac{1}{2m}$ hatten wir (siehe Vorlesung):

$$n = \frac{1}{V} \frac{z}{1-z} + \frac{g_{3/2}(z)}{\lambda^3}$$



D.h. für $T > T_c \Rightarrow z < 1$, der Grundzustandsterm verschwindet für $V \rightarrow \infty$
 $T < T_c \Rightarrow z = 1$, $\frac{1}{V} \frac{z}{1-z} = \frac{N_0}{V}$ bleibt endlich.

Kritische Bedingung: $n = \frac{g_{3/2}(1)}{\lambda^3} = \frac{\zeta(3/2)}{\lambda^3}$

Was passiert im allg. Fall? Untersuche $g_{d/r}(1) = \zeta(d/r)$:

$$g_\alpha(1) = \frac{1}{\Gamma(\alpha)} \int_0^\infty dx \frac{x^{\alpha-1}}{e^x - 1}$$

\Rightarrow keine Probleme für $x \rightarrow \infty$,
 aber für $x \rightarrow 0$: $e^x \approx 1 + x + \dots$

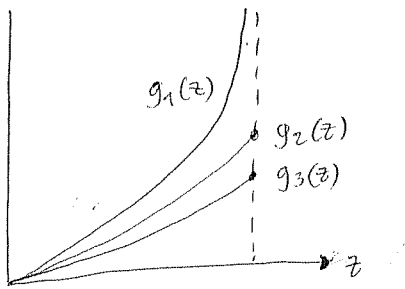
\Rightarrow Integrand: $x^{\alpha-2}$

\Rightarrow Integral konvergiert nur für $\alpha-2 > -1$,
 d.h. $\alpha > 1$

$\Rightarrow \zeta(\alpha \leq 1)$ divergiert.

\Rightarrow dann hat n immer Schnittpunkt für $z < 1$,
 keine BEC.

Dh $d/r > 1$, damit Bose-Einstein-Kondensation eintreten kann. Ansonsten divergiert $g_{d/r}(1)$, und $n < \dots g_{d/r}(1) = \infty$ ist immer erfüllt.



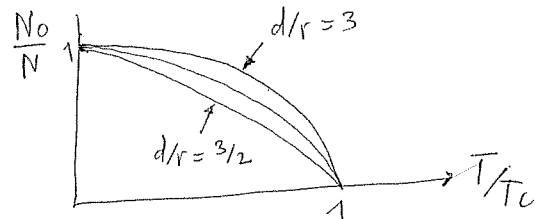
$d=3, r=2 \rightarrow$ BEC (nichtrelativistisch)
 $d=3, r=1 \rightarrow$ BEC (ultrarelativistisch)
 $d=2, r=2 \rightarrow$ keine BEC! (nichtrel.)
 $d=2, r=1 \rightarrow$ BEC

d) $N_0(T) = ?$

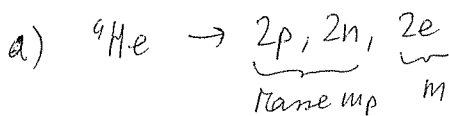
$$\frac{N}{V} = \frac{N_0}{V} + (\dots) g_{d/r}(z); \text{ im Kondensationsgebiet: } z=1$$

$$\Rightarrow \frac{N_0}{N} = 1 - \frac{S_d \Gamma(d/r) \zeta(d/r)}{n r h^d a^{d/r}} (kT)^{d/r} = 1 - \left(\frac{kT}{kT_c} \right)^{d/r}$$

$$\Rightarrow kT_c = \left[\frac{n r h^d a^{d/r}}{S_d \Gamma(d/r) \zeta(d/r)} \right]^{r/d}$$



29) Weisses Zwerg



\Rightarrow Insgesamt N Elektronen, N Protonen, N Neutronen:

$$M = (m + 2m_p) N \approx 2m_p N$$

$$V = \frac{4\pi}{3} R^3$$

$$\Rightarrow n = \frac{N}{V} = \frac{3M}{8\pi m_p R^3}$$

Impulskugel:

$$N = 2 \underbrace{\frac{V}{h^3}}_{\text{Spin}} \int_{\text{PF}} d^3p = \frac{2V}{(2\pi\hbar)^3} \frac{4\pi}{3} p_F^3 = \frac{V}{3\pi^2 \hbar^3} p_F^3 \Rightarrow p_F = \hbar (3\pi^2 n)^{1/3}$$

$$\Rightarrow x_F = \frac{p_F}{mc} = \frac{\hbar}{mc} (3\pi^2 n)^{1/3} = \frac{\hbar}{mc} \left(\frac{9\pi M}{8m_p R^3} \right)^{1/3} = \frac{\hbar}{mc} \left(\frac{9\pi}{8m_p} \right)^{1/3} \frac{M^{1/3}}{R}$$

und: $\sum_p \rightarrow \frac{V}{\pi^2 \hbar^3} \int dp p^2$

b) Energie:

1 Teilchen: $E_p = \sqrt{m^2 c^4 + p^2 c^2}$

Gesamt: $E_c = \int dE g(E) \underbrace{f(E)}_{\approx 1, \text{ da } T \approx 0} E$

ben:

$$E_c = \sum_p E_p \approx \frac{V}{\pi^2 \hbar^3} \int_0^{p_F} dp p^2 \sqrt{m^2 c^4 + p^2 c^2} = \frac{V}{\pi^2} \frac{m c^2}{\hbar^3} \int_0^{p_F} dp p^2 \sqrt{1 + \underbrace{\frac{p^2}{m^2 c^2}}_{x^2}}$$

$$= \frac{V}{\pi^2} \left(\frac{m c}{\hbar} \right)^3 m c^2 \int_0^{x_F} dx x^2 \sqrt{1+x^2}$$

c) Nichtrelativistisch:

$$x_F \ll 1 \Rightarrow \int_0^{x_F} dx x^2 \sqrt{1+x^2} \approx \int_0^{x_F} dx x^2 \left(1 + \frac{x^2}{2} + \dots \right) = \frac{x_F^3}{3} + \frac{x_F^5}{10} + \dots$$

$$\frac{x_F^3}{3} = \left(\frac{\hbar}{m c} \right)^3 \frac{(3\pi^2 n)}{3} \Rightarrow E_c = N m c^2 \checkmark$$

Ultrarelativistisch:

$$x_F \gg 1 \Rightarrow \int_0^{x_F} dx x^2 \sqrt{1+x^2} = \int_0^{x_F} dx x^3 \sqrt{1+\frac{1}{x^2}} \approx \int_0^{x_F} dx x^3 \left(1 + \frac{1}{2x^2} + \dots \right) = \frac{x_F^4}{4} + \frac{x_F^2}{4} + \dots$$

$$\Rightarrow E_c = \overset{\sim R^3}{\underbrace{\frac{V}{4\pi^2}}_{\substack{\text{Integral f\"ur} \\ \text{kleine } x \text{ klein}}}} \left(\frac{m c}{\hbar} \right)^3 m c^2 x_F^4 \left(1 + \frac{1}{x_F^2} \right)$$

$$= \underbrace{\frac{\hbar c}{3\pi} \left(\frac{g\pi^2}{8m_p} \right)^{4/3}}_{C_1} \frac{1}{R} \left[1 + \underbrace{\left(\frac{m c}{\hbar} \right)^2 \left(\frac{8m_p}{g\pi^2} \right)^{2/3}}_{C_2} R^2 + \dots \right]$$

d) Nullpunktsdruck, $x_F \gg 1$:

$$P = - \frac{\partial E_c}{\partial V} = - \frac{1}{4\pi^2} \left(\frac{m c}{\hbar} \right)^3 m c^2 \left[x_F^4 + \underbrace{V \frac{\partial}{\partial V} x_F^4} \right]$$

$$= \frac{1}{12\pi^2} \left(\frac{m c}{\hbar} \right)^3 m c^2 x_F^4 \quad \underbrace{V \cdot 4 x_F^3 \frac{\partial x_F}{\partial n} \frac{\partial n}{\partial V}}_{\frac{x_F}{3n} - \frac{N}{V^2} = -\frac{n}{V}} = -\frac{4}{3} x_F^2$$

$$= \frac{\hbar c}{12\pi^2} (3\pi^2 n)^{4/3}$$

$$e) \quad E = \underbrace{\frac{c_1}{R} + c_1 c_2 R}_{\text{Elektronen}} - \underbrace{\frac{3}{5} \frac{GM^2}{R}}_{\text{He-Kerne}} = \underbrace{\left(c_1 - \frac{3}{5} GM^2\right)}_{>0} \frac{1}{R} + c_1 c_2 R$$

d.h. Minimum nur für

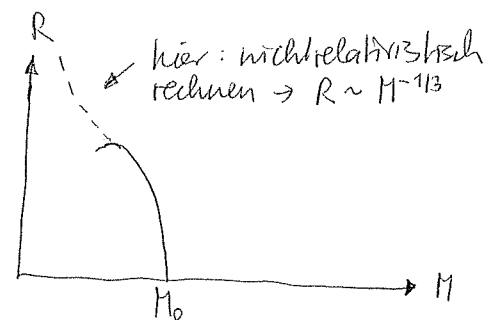
$$c_1 > \frac{3}{5} GM^2 \Rightarrow \frac{5}{3G} \frac{\hbar c}{3\pi} \left(\frac{9\pi}{8m_p}\right)^{4/3} > M^{2/3}$$

$$\Rightarrow M < M_0 = \left[\frac{\hbar c}{3\pi} \left(\frac{9\pi}{8m_p}\right)^{4/3} \frac{5}{3G} \right]^{3/2} = 1.73 M_\odot, \text{ mit } M_\odot \approx 2 \cdot 10^{30} \text{ kg}$$

$$\underline{R_0(M)}: \quad \frac{\partial E}{\partial R} = -\frac{1}{R^2} \left(c_1 - \frac{3}{5} GM^2\right) + c_1 c_2 = 0$$

$$\Rightarrow R^2 = \frac{1}{c_2} \left(1 - \frac{3GM^2}{5c_1}\right) = \frac{1}{c_2} \left(1 - \left(\frac{M}{M_0}\right)^{2/3}\right)$$

$$\Rightarrow R = \frac{\hbar}{mc} \left(\frac{9\pi}{8m_p}\right)^{1/3} M^{1/3} \sqrt{1 - \left(\frac{M}{M_0}\right)^{2/3}}$$



By the way: typische Grössenordnungen:

$$x_F \sim 2$$

$$E_F = mc^2 \sqrt{1+x_F^2} \sim 1 \text{ MeV}$$

$$T = 10^7 \text{ K} \Rightarrow kT \sim 1 \text{ keV} \ll E_F$$

$$M \sim M_\odot$$

$$R \sim 5000 \text{ km}$$